

Documentation

Dataset for Manuscript "Fast Simulation of Wide-Angle Coherent Diffractive Imaging"

Paul Tuemmler, Julia Apportin, Thomas Fennel, Christian Peltz

1 General Information

Dataset title:	Dataset for Manuscript "Fast Simulation of Wide-Angle Coherent Diffractive Imaging"
Authors/Creators:	Paul Tuemmler, Julia Apportin, Thomas Fennel, Christian Peltz
Affiliation:	Institute of Physics, University of Rostock, 18051 Rostock, Germany
E-Mail:	christian.peltz@uni-rostock.de
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2 Description

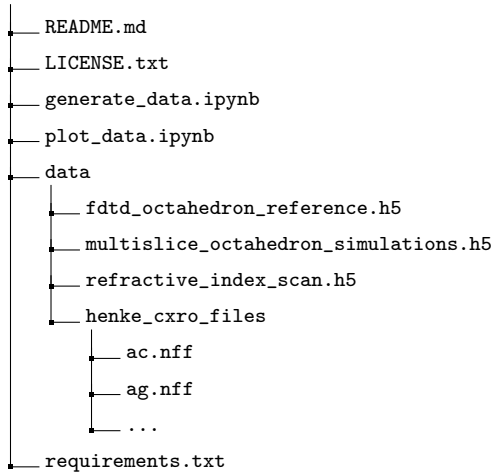
This archive contains the data shown in publication [1] together with a python script to visualize them. It also provides a python script to generate this data using the open source implementation of pMSFT.

3 Data Format

The data is stored in a compressed HDF5 format, which can be loaded using virtually all programming languages. To view the data without extensive programming knowledge, we recommend using `silx view` (<https://www.silx.org/doc/silx/latest/applications/view.html>) or `HDFView` (<https://www.hdfgroup.org/downloads/hdfview/>).

4 Archive Structure

The archive has the following structure:



The contents of the individual files and folders are described below in more detail.

4.1 fdttd_octahedron_reference.h5

Compressed HDF5 file containing the reference FDTD simulation data for the benchmark example shown in Figure 1 of the publication.

```
fdtd_octahedron_reference.h5
├── fdttd_image : 2d array of the FDTD reference image (256x256 pixels; float64)
├── fdttd_kx : kx grid corresponding to fdttd_image (256x256 pixels; float64)
├── fdttd_ky : ky grid corresponding to fdttd_image (256x256 pixels; float64)
├── fdttd_kz : kz grid corresponding to fdttd_image (256x256 pixels; float64)
├── model : 3d array of the sample model (150x150x150 voxels; uint8)
├── model_width_in_nm : box width of the sample model in nm (float64)
├── wavelength_in_nm : wavelength used in the simulation in nm (float64)
└── refractive_index : complex refractive index used in the simulation (complex128)
```

4.2 multislice_octahedron_simulations.h5

Compressed HDF5 file containing the simulation results for the comparison of pMSFT and the other methods with the FDTD reference data (see above) for an octahedron as shown in Figure 1 of the publication. This can also be generated using `generate_data.ipynb`, see below for more details.

```
multislice_octahedron_simulations.h5
├── images
│   ├── BORN : 2d array of the Born's approximation result (2048x2048 pixels; float64)
│   ├── HARE : 2d array of result using Hare's method (2048x2048 pixels; float64)
│   └── MSFT : 2d array of the MSFT result (2048x2048 pixels; float64)
```

- PMSFT : 2d array of the pMSFT result (2048x2048 pixels; float64)
 - SAXS : 2d array of the SAXS result (2048x2048 pixels; float64)
 - pmsft_kx : kx grid corresponding to the above images (2048x2048 pixels; float64)
 - pmsft_ky : ky grid corresponding to the above images (2048x2048 pixels; float64)
 - pol_mask : linear polarization mask (2048x2048 pixels; float64)

4.3 refractive_index_scan.h5

Compressed HDF5 file containing the simulation results for the refractive index scan shown in Figure 3 of the publication.

```

refractive_index_scan.h5
├── real_space : axis corresponding to real parts of the refractive index (512 pixels; float64)
├── imag_space : axis corresponding to imaginary parts of the refractive index (256 pixels; float64)
├── difference_map_Q
│   ├── BORN : 2d array of Q value in Born's approximation based on the above grid (256x512 pixels; float64)
│   ├── HARE : 2d array of Q value using Hare's method based on the above grid (256x512 pixels; float64)
│   ├── MSFT : 2d array of the MSFT Q value based on the above grid (256x512 pixels; float64)
│   └── PMSFT : 2d array of the pMSFT Q value based on the above grid (256x512 pixels; float64)
├── difference_map_R
│   ├── BORN : 2d array of R value in Born's approximation based on the above grid (256x512 pixels; float64)
│   ├── HARE : 2d array of R value using Hare's method based on the above grid (256x512 pixels; float64)
│   ├── MSFT : 2d array of the MSFT R value based on the above grid (256x512 pixels; float64)
│   └── PMSFT : 2d array of the pMSFT R value based on the above grid (256x512 pixels; float64)
├── difference_map_R_mask
│   ├── BORN : 2d array of R value (with mask) in Born's approximation based on the above grid (256x512 pixels; float64)
│   ├── HARE : 2d array of R value (with mask) using Hare's method based on the above grid (256x512 pixels; float64)
│   ├── MSFT : 2d array of the MSFT R value (with mask) based on the above grid (256x512 pixels; float64)
│   └── PMSFT : 2d array of the pMSFT R value (with mask) based on the above grid (256x512 pixels; float64)
├── highlighted_points
│   └── 0
│       ├── ref_index : refractive index used for this point (complex128)
│       ├── image_BORN : 2d array of the Born's approximation result for point 0 (2048x2048 pixels; float64)
│       ├── image_HARE : 2d array of result using Hare's method for point 0 (2048x2048 pixels; float64)
│       ├── image_MSFT : 2d array of the MSFT result for point 0 (2048x2048 pixels; float64)
│       ├── image_PMSFT : 2d array of the pMSFT result for point 0 (2048x2048 pixels; float64)
│       └── mie_perp : 1d array of the Mie result in perpendicular direction (2048 pixels; float64)

```

```

└─mie_polar : 1d array of the Mie result in polarization direction (2048 pixels; float64)
└─ 1 ...
└─ 2 ...
└─ kx : kx grid corresponding to the above images (2048x2048 pixels; float64)
└─ ky : ky grid corresponding to the above images (2048x2048 pixels; float64)
└─ angles : grid of the angles corresponding to the images (2048x2048 pixels; float64)
└─ pol_mask : linear polarization mask (2048x2048 pixels; float64)
└─ max_angle : maximum considered scattering angle in degrees (float64)
└─ wavelength : wavelength used in the simulation in nm (float64)
└─ Rsphere : radius of the spherical sample in nm (float64)

```

4.4 Other files

- `data/henke_cxro_files/`:
Folder containing tabulated complex refractive index data from CXRO Henke database (https://henke.lbl.gov/optical_constants/asf.html) (state as of 30.09.2025), used for Figure 3 of the publication.
- `generate_data.ipynb`:
Jupyter notebook containing the source code to generate the `refractive_index_scan.h5` and `multislice_octahedron_simulations.h5` data files.
- `plot_data.ipynb`:
Jupyter notebook containing the source code to generate the plots shown in the publication using the provided data files.

4.5 Using the Jupyter Notebooks

We strongly recommend setting up a virtual environment for the execution of these scripts, i.e. using conda (<https://docs.conda.io/en/latest/>). This code has been tested for python 3.12. Create and activate a conda environment using

```
conda create -n pymsft python=3.12
conda activate pymsft
```

At this point you can install the package requirements using pip. If you want to calculate the data yourself you should install the full requirements using

```
pip install -r requirements.txt
```

If, however, you are only interested in plotting the data, you can instead use a reduced set of required packages using

```
pip install -r requirements_plotting.txt
```

To take advantage of GPU acceleration for the calculation scripts, we are using CuPy (<https://cupy.dev/>). See the installation instructions (<https://docs.cupy.dev/en/stable/install.html>) for more details. The simplest way to install it is currently via conda:

```
conda install -c conda-forge cupy==13.6.0
```

Note that an NVIDIA GPU with the corresponding driver installed is required.

References

- [1] Paul Tuemmler, Julia Apportin, Thomas Fennel and Christian Peltz, *Fast Simulation of Wide-Angle Coherent Diffractive Imaging*, Laser & Photonics Reviews (<https://doi.org/10.1002/lpor.202502001>).